

Cosmic Microwave Background Radiation Foregrounds

Foreground: study the radiation from the Milky Way at 5GHz and at 10GHz on both Hemispheres (Portugal and Brazil) for space data processing.

THE 5GHz EXPERIMENT





Tsys < 20 K; B = 200 MHz; 104 dB gain
High-performance IF strip
Latest RF tech + microstrip design + MMIC
Zero-IF Converter + I,Q modulation
Digital Correlator : 4-channel, FPGA implemented processing 16Gbps!
An SKA (Potential!) Digital Demodulator

Dynamic Range: Total=20dB, Instantaneous=80dB

Suitable for state of the art RA applications.

MoU with ESA Planck Science Team.



For this project was implemented a C-band (5GHz) Low Noise Amplifier (LNA) using new low noise Pseudomorphic High Electron Mobility Transistors (pHEMTS) from Avago. The amplifier was developed to be a cost effective solution in a receiver chain for Galactic Emission Mapping (GEM-P) project in Portugal.

Generational Conference on Computer as a Tool (EUROCON), 2011 IEEE , vol., no., pp.1-4, 27-29 April 2011 10: 10.1109/EUROCON.2011.5929355

GEM-Portugal Antenna





The receiver system is equipped with a novel receiver with a full digital back-end using a low-cost Field Programmable Gate Array without compromising its performance relation. This new digital backend comprises a base-band complex cross-correlator outputting the four Stokes parameters of the incoming

Bergano, J.B.; F. F. Fernandes; L. Cupido; Barbosa, D.; R. Fonseca; I. S. V. Ferreira; B. Grossan; G. F. S. Smoot; "Digital Complex Correlator for a C-band Polarimetry survey", *Experimental Astronomy*, Vol. 30, No. 1, pp. 23 - 37, February, 2011.

GEM-Brasil Antenna

THE 10 GHz EXPERIMENT

Development of a full digital polarimeter;

• Measurement of the Stokes-Parameters (circular polarization)

• Measure the change (ΔT) in the antenna equivalent noise temperature (TA) which is extremely low (10K). Targets:

- Sensitivity (ΔT) is 0.2 mK
- Equivalent Temperature of the Receiver is 17 K
- Center Frequency = 10GHz with 1GHz
 Bandwidth
- Technology comparison:
- Low Noise Front-end →LNA
- Cost-effective performing solution



- Measuring Cosmic Noise Signal Signatures
- Research topic very important in radio-astronomy
 Equipment complex and expensive
- Several Architectures reviewed a

Our Proposal: a cheap but potentially performing solution

- Total Power approach with very large bandwidth
 - Low Noise front –end development
 Temperature stabilized IF circuitry
- Implementation of an innovative digital detector based on high data rate sampling and FPGA
 - processing
- Approach relevant to other niches:
- Telecommunications and Earth-Sensing
- Versatile solution: integration times, resolution, calibration, easy reproduction of the hardware, etc



